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Chiral three-nucleon force

The three-nucleon force problem Chiral EFT for nuclear forces Chiral expansion of the nuclear force up to N⁴LO Inclusion of the Δ(1232) isobar Summary & outlook



The three-nucleon force problem

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PHYSICAL REVIEW

VOLUME 55

Many-Body Interactions in Atomic and Nuclear Systems

H. PRIMAKOFF, Polytechnic Institute of Brooklyn, Brooklyn, New York

AN D

T. HOLSTEIN,* New York University, University Heights, New York, New York (Received March 28, 1938)

"...replacement of field interactions by two-body action-at-a-distance potentials is a poor approximation in nuclear physics."

The three-nucleon force problem

Modern phenomenological NN potentials (AV18, CDBonn, Nijm I,II, Reid93, ...)

- Long-range part due to EM interaction and the one-pion exchange potential
- Short-range pieces modeled phenomenological; benefit from the general form of the NN potential being rather simple:

 $V(\vec{p}', \vec{p}) \equiv \langle \vec{p}' | V | \vec{p} \rangle = \left\{ \mathbf{1}_{\text{spin}}, \ \vec{\sigma}_1 \cdot \vec{\sigma}_2, \ S_{12}(\vec{q}), \ S_{12}(\vec{k}), \ i\vec{S} \cdot \vec{q} \times \vec{k}, \ \vec{\sigma}_1 \cdot \vec{q} \times \vec{k} \vec{\sigma}_2 \cdot \vec{q} \times \vec{k} \right\} \times \{ \mathbf{1}_{\text{isospin}}, \ \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2 \}$ • Perfect fit to NN data ($\chi^2 \sim 1$)



Chiral expansion of nuclear forces

- Triton binding energy calculated based on V_{NN} is typically under-bound by the amount of ~ 1 MeV
- Large discrepancies for the total nucleon-deuteron cross section

→ "evidence" for missing 3NFs

Phenomenological 3NF models FM, Brasil, TM, Urbana, Illinois,...





Chiral expansion of nuclear forces











Fugita-Miyazawa, Tucson-Melbourne, Brasil, Urbana IX, Illinois, ...

0



A_y (d)

exp-theory (NN+3NF) 0 10 70

A_y (**d**)

Fugita-Miyazawa Tucson-Melbourne, Brasil, Urbana IX, Illinois, ...

0



from: Kalantar-Nayestanaki, EE, Messchendorp, Nogga, Rept. Prog. Phys. 75 (2012) 016301

General structure of a 3NF

Why is it so difficult to model the 3NF as compared to NN potentials?

- More scarce Nd data base compared to np and pp data bases
- Solving the Faddeev equation for 3N more involved than solving the LS equation for NN
- General structure of the 3NF is much more involved

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Most general structure of a local, isospin-symmetric 3NF

Krebs, Gasparyan, EE '13; Phillips, Schat '13; EE, Gasparyan, Krebs, Schat, in preparation

Generators \mathcal{G} in momentum space	Generators $\tilde{\mathcal{G}}$ in coordinate space
$\mathcal{G}_1 = 1$	$\tilde{\mathcal{G}}_1 = 1$
$\mathcal{G}_2 = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3$	$ ilde{\mathcal{G}}_2 = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3$
$\mathcal{G}_3=ec{\sigma}_1\cdotec{\sigma}_3$	$ ilde{\mathcal{G}}_3=ec{\sigma}_1\cdotec{\sigma}_3$
$\mathcal{G}_4 = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3 ec{\sigma}_1 \cdot ec{\sigma}_3$	$ ilde{\mathcal{G}}_4 = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3 ec{\sigma}_1 \cdot ec{\sigma}_3$
$\mathcal{G}_5 = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 ec{\sigma}_1 \cdot ec{\sigma}_2$	$ ilde{\mathcal{G}}_5 = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 ec{\sigma}_1 \cdot ec{\sigma}_2$
$\mathcal{G}_6 = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_1 \cdot (ec{\sigma}_2 imes ec{\sigma}_3)$	$ ilde{\mathcal{G}}_6 = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_1 \cdot (ec{\sigma}_2 imes ec{\sigma}_3)$
$\mathcal{G}_7 = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_2 \cdot (ec{q}_1 imes ec{q}_3)$	$ ilde{\mathcal{G}}_7 = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_2 \cdot (\hat{r}_{12} imes \hat{r}_{23})$
$\mathcal{G}_8 = ec{q_1} \cdot ec{\sigma}_1 ec{q_1} \cdot ec{\sigma}_3$	$ ilde{\mathcal{G}}_8 = \hat{r}_{23} \cdot ec{\sigma}_1 \hat{r}_{23} \cdot ec{\sigma}_3$
$\mathcal{G}_9 = ec{q_1} \cdot ec{\sigma}_3 ec{q_3} \cdot ec{\sigma}_1$	$ ilde{\mathcal{G}}_9 = \hat{r}_{23} \cdot ec{\sigma}_3 \hat{r}_{12} \cdot ec{\sigma}_1$
${\cal G}_{10}=ec q_1\cdotec \sigma_1ec q_3\cdotec \sigma_3$	$ ilde{\mathcal{G}}_{10} = \hat{r}_{23} \cdot ec{\sigma}_1 \hat{r}_{12} \cdot ec{\sigma}_3$
$\mathcal{G}_{11} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 ec{q}_1 \cdot ec{\sigma}_1 ec{q}_1 \cdot ec{\sigma}_2$	$ ilde{\mathcal{G}}_{11} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 \hat{r}_{23} \cdot ec{\sigma}_1 \hat{r}_{23} \cdot ec{\sigma}_2$
$\mathcal{G}_{12} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 ec{q}_1 \cdot ec{\sigma}_1 ec{q}_3 \cdot ec{\sigma}_2$	$ ilde{\mathcal{G}}_{12} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 \hat{r}_{23} \cdot ec{\sigma}_1 \hat{r}_{12} \cdot ec{\sigma}_2$
$\mathcal{G}_{13} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 ec{\sigma}_1 ec{q}_1 \cdot ec{\sigma}_2$	$ ilde{\mathcal{G}}_{13} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 \hat{r}_{12} \cdot ec{\sigma}_1 \hat{r}_{23} \cdot ec{\sigma}_2$
$\mathcal{G}_{14} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 ec{\sigma}_1 ec{q}_3 \cdot ec{\sigma}_2 ec{\sigma}_2$	$ ilde{\mathcal{G}}_{14} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 \hat{r}_{12} \cdot ec{\sigma}_1 \hat{r}_{12} \cdot ec{\sigma}_2$
$\mathcal{G}_{15} = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3 ec{q}_2 \cdot ec{\sigma}_1 ec{q}_2 \cdot ec{\sigma}_3$	$ ilde{\mathcal{G}}_{15} = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3 \hat{r}_{13} \cdot ec{\sigma}_1 \hat{r}_{13} \cdot ec{\sigma}_3$
$\mathcal{G}_{16} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 ec{q}_3 \cdot ec{\sigma}_2 ec{q}_3 \cdot ec{\sigma}_3$	$ ilde{\mathcal{G}}_{16} = oldsymbol{ au}_2 \cdot oldsymbol{ au}_3 \hat{r}_{12} \cdot ec{\sigma}_2 \hat{r}_{12} \cdot ec{\sigma}_3$
$\mathcal{G}_{17} = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3 ec{q}_1 \cdot ec{\sigma}_1 ec{q}_3 \cdot ec{\sigma}_3$	$ ilde{\mathcal{G}}_{17} = oldsymbol{ au}_1 \cdot oldsymbol{ au}_3 \hat{r}_{23} \cdot ec{\sigma}_1 \hat{r}_{12} \cdot ec{\sigma}_3$
${\cal G}_{18} = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_1 \cdot ec{\sigma}_3 ec{\sigma}_2 \cdot (ec{q}_1 imes ec{q}_3)$	$ ilde{\mathcal{G}}_{18} = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_1 \cdot ec{\sigma}_3 ec{\sigma}_2 \cdot (\hat{r}_{12} imes \hat{r}_{23})$
${\mathcal{G}}_{19} = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_3 \cdot ec{q}_1 ec{q}_1 \cdot (ec{\sigma}_1 imes ec{\sigma}_2)$	$ ilde{\mathcal{G}}_{19} = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_3 \cdot \dot{r}_{23} \dot{r}_{23} \cdot (ec{\sigma}_1 imes ec{\sigma}_2)$
$\mathcal{G}_{20} = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_1 \cdot ec{q}_1 ec{\sigma}_3 \cdot ec{q}_1 ec{\sigma}_2 \cdot (ec{q}_1 imes ec{q}_3)$	$ ilde{\mathcal{G}}_{20} = oldsymbol{ au}_1 \cdot (oldsymbol{ au}_2 imes oldsymbol{ au}_3) ec{\sigma}_1 \cdot \hat{r}_{23} ec{\sigma}_3 \cdot \hat{r}_{12} ec{\sigma}_2 \cdot (\hat{r}_{12} imes \hat{r}_{23})$



Assuming hermiticity, time reversal & parity invariance, **20 structure functions** are needed:

$$V(r_{12}, r_{23}, r_{31}) = \sum_{i=1}^{20} \tilde{\mathcal{G}}_i F_i(r_{12}, r_{23}, r_{31})$$

+ permutations

+ permutations

$$V(q_1, q_2, q_3) = \sum_{i=1}^{20} \mathcal{G}_i F_i(q_1, q_2, q_3)$$

+ permutations

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Phenomenological modeling seems not feasible!

Need a theoretical approach which would:
be based on QCD,
yield consistent many-body forces,
be systematically improvable,
allow for error estimation

→ Chiral Effective Field Theory



Chiral dynamics and the pion-nucleon system

Chiral Perturbation Theory: expansion of the scattering amplitude in powers of Weinberg, Gasser, Leutwyler, Meißner, ...

 $Q = \frac{\text{momenta of pions and nucleons or } M_{\pi} \sim 140 \text{ MeV}}{\text{hard scales [at best } \Lambda_{\chi} = 4\pi F_{\pi} \sim 1 \text{ GeV}]}_{\text{Manohar, Georgi '84}}$

Tool: Feynman calculus using the effective chiral Lagrangian



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Tool: Feynman calculus using the effective chiral Lagrangian



Pion-nucleon scattering up to Q⁴ in heavy-baryon ChPT

Fettes, Meißner '00; Krebs, Gasparyan, EE '12



Improving the convergence of ChPT:

- Covariant formulations (no 1/m_N-expansion) Becher, Leutwyler, Fuchs, Gegelia, Japaridze, Scherer ...
- Explicit treatment of the Δ(1232) isobar Jenkens, Manohar, Hemmert, Holstein, Kambor, ...

For ChPT to be useful, (renormalized) LECs must be natural, i.e. $\sim \alpha_i / \Lambda_x^n$, $\alpha_i = \mathcal{O}(1)$



$$c_2^\Delta = -c_3^\Delta = 2c_4^\Delta = rac{4h_A^2}{9(m_\Delta-m_N)} \simeq 2.8\,{
m GeV}^{-1}$$
 Bernard, Kaiser, Meißner '97

 $ar{d}_{14}^{\Delta} - ar{d}_{15}^{\Delta} = -2(ar{d}_1^{\Delta} + ar{d}_2^{\Delta}) = 2ar{d}_3^{\Delta} = rac{-2h_A^2}{9(m_{\Delta} - m_N)^2} \simeq -4.8\,\mathrm{GeV}^{-2}$ Krebs, Gasparyan, EE, to appear

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LECs from pion-nucleon scattering (HB ChPT) in units of GeV ⁻ⁿ Krebs, Gasparyan, EE, to appear; similar results found by Fettes, Meißner; Büttiker, Meißner,													
	c_1	c_2	c_3	c_4	$\bar{d}_1 + \bar{d}_2$	\bar{d}_3	\bar{d}_5	$\bar{d}_{14} - \bar{d}_{15}$	\bar{e}_{14}	\bar{e}_{15}	\bar{e}_{16}	\bar{e}_{17}	\bar{e}_{18}
Δ-less approach	-0.75	3.49	-4.77	3.34	6.21	-6.83	0.78	-12.02	1.52	-10.41	6.08	-0.37	3.26
∆-full approach	-0.95	1.90	-1.78	1.50	2.40	-3.87	1.21	-5.25	-0.24	-6.35	2.34	-0.39	2.81
Δ-contribution	0	2.81	-2.81	1.40	2.39	-2.39	0	-4.77	1.87	-4.15	4.15	-0.17	1.32

The hope: LECs of a more natural size \longrightarrow better convergence of the EFT expansion...

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ChPT, DOF: π , N

Expansion in $q \in \left(\frac{M_{\pi}}{\Lambda_{\chi}}, \frac{p_i}{\Lambda_{\chi}}\right)$

simple, smaller number of LECs

certain LECs are unnaturally large
 c₂ = −2.8, c₃ = −3.9, c₄ = 2.9 [GeV⁻¹]
 → (sometimes) slow convergence...

ChPT mit Δ , **DOF**: π , N, Δ Expansion in: $\epsilon \in \left(\frac{M_{\pi}}{\Lambda}, \frac{p_i}{\Lambda}, \frac{m_{\Delta} - m_N}{\Lambda}\right)$ more LECs, considerably more extensive $S^{\mu\nu}(p) = \frac{p + m_{\Delta}}{p^2 - m_{\Delta}^2} \left(-g^{\mu\nu} + \frac{1}{3} \gamma^{\mu} \gamma^{\nu} + \frac{1}{3m_{\Delta}} (\gamma^{\mu} p^{\nu} - \gamma^{\nu} p^{\mu}) + \frac{2}{3m_{\Delta}^2} p^{\mu} p^{\nu} \right)$ more natural LECs, e.g. $c_2 = -0.3, c_3 = -0.8, c_4 = 1.3 \,[\text{GeV}^{-1}]$

Siemens, Bernard, EE, Krebs, Meißner '14



Siemens, Bernard, EE, Krebs, Meißner '14



Covariant ChPT



Siemens, Bernard, EE, Krebs, Meißner '14







Siemens, Bernard, EE, Krebs, Meißner '14



Covariant ChPT $\pi^- p \rightarrow \pi^0 \pi^0 n$ $\pi^- p \rightarrow \pi^+ \pi^- n$ 1.4 3.0 1.2 2.5 1.0 2.0 ຊີ ມີ 0.6 1.5 1.0 0.4 0.5 0.2 0.0 0.0 0.25 0.30 0.35 0.40 0.30 0.20 0.20 0.25 0.35 0.40 $T_{\pi}[\text{GeV}]$ $T_{\pi}[\text{GeV}]$





Chiral dynamics and nuclear forces









A new, soft scale associated with nuclear binding

 $Q \sim 1/a_S \simeq 8.5 \text{ MeV}(36 \text{ MeV})$ in ${}^1S_0 ({}^3S_1)$ to be generated dynamically (need resummations...)



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Pionless EFT (valid for $\sqrt{m_N E_B} \ll Q \ll M_{\pi}$)

- zero-range forces between nucleons
- for 2N equivalent to Effective Range Theory
- universality, Efimov physics, cold gases, halos,...



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<u>**Chiral EFT**</u> (valid for $Q \sim M_{\pi}$)

Weinberg, van Kolck, EE, Glöckle, Meißner, Machleidt, Entem...

- Schrödinger equation for nucleons interacting via contact forces and long-range potentials (pion exchanges)



$$\left[\left(\sum_{i=1}^{A} \frac{-\vec{\nabla}_{i}^{2}}{2m_{N}} + \mathcal{O}(m_{N}^{-3})\right) + \underbrace{V_{2N} + V_{3N} + V_{4N} + \dots}_{\text{derived in ChPT}}\right] |\Psi\rangle = E|\Psi\rangle$$

- access to heavier nuclei (ab initio few-/many-body methods)

Chiral expansion of nuclear forces



2N force: accurate N³LO potentials are available Entem-Machleidt '03; EE-Glöckle-Meißner '04

3N force: N²LO 3NF included in most calculations N³LO 3NF worked out Bernard, EE, Krebs Meißner '08,'11; (probably) not yet converged → higher orders numerical PWD developed Golak, Skibinski, Krebs, Hebeler, ..., first results available Witala et al.'13

4N force: leading (i.e. N³LO) terms worked out EE '06; contrib. to ⁴He BE ~ few 100 keV Rospedzik et al. '06

Chiral expansion of nuclear forces



The "standard" nuclear chiral Hamiltonian has been extensively tested in few- and many-body systems

Nucleon-nucleon force up to N³LO

Ordonez et al. '94; Friar & Coon '94; Kaiser et al. '97; E.E. et al. '98, '03; Kaiser '99-'01; Higa, Robilotta '03; ...



van Kolck et al. '93, '96; Friar et al. '99 '03, '04; Niskanen '02; Kaiser '06; E.E. et al. '04,'05,'07; ...

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E.E. et al. '04,'05,'07; ...

by chiral symmetry and exp. information on the πN system

Chiral 2π exchange (upto N²LO)

 $\mathcal{V}_{NN} \;=\; V_C(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_C(r) \;+\; \left[V_S(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_S(r) \right] \vec{\sigma}_1 \cdot \vec{\sigma}_2$

 $+ \left[V_T(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_T(r) \right] \left(3 \vec{\sigma}_1 \cdot \hat{r} \vec{\sigma}_2 \cdot \hat{r} - \vec{\sigma}_1 \cdot \vec{\sigma}_2 \right) + \left[V_{LS}(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_{LS}(r) \right] \vec{L} \cdot \vec{S} \,,$

Chiral 2π exchange (upto N²LO)

$$\begin{aligned} \mathcal{V}_{NN} &= V_C(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_C(r) + \left[V_S(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_S(r) \right] \vec{\sigma}_1 \cdot \vec{\sigma}_2 \\ &+ \left[V_T(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_T(r) \right] \left(3\vec{\sigma}_1 \cdot \hat{r}\vec{\sigma}_2 \cdot \hat{r} - \vec{\sigma}_1 \cdot \vec{\sigma}_2 \right) + \left[V_{LS}(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_{LS}(r) \right] \vec{L} \cdot \vec{S} \,, \end{aligned}$$

The profile functions (in Dimensional Regularization)

$$\begin{split} V_C^{TPE}(r) &= \frac{3g^2m^6}{32\pi^2f^4} \frac{e^{-2x}}{x^6} \Big\{ \left(2c_1 + \frac{3g^2}{16M} \right) x^2 (1+x)^2 + \frac{g^5x^5}{32M} + \left(c_3 + \frac{3g^2}{16M} \right) \left(6 + 12x + 10x^2 + 4x^3 + x^4 \right) \\ W_T^{TPE}(r) &= \frac{g^2m^6}{48\pi^2f^4} \frac{e^{-2x}}{x^6} \Big\{ - \left(c_4 + \frac{1}{4M} \right) (1+x)(3+3x+x^2) + \frac{g^2}{32M} \left(36 + 72x + 52x^2 + 17x^3 + 2x^4 \right) \Big\}, \\ V_T^{TPE}(r) &= \frac{g^4m^5}{128\pi^3f^4x^4} \Big\{ - 12K_0(2x) - (15+4x^2)K_1(2x) + \frac{3\pi m e^{-2x}}{8Mx} \left(12x^{-1} + 24 + 20x + 9x^2 + 2x^3 \right) \Big\}, \\ W_C^{TPE}(r) &= \frac{g^4m^5}{128\pi^3f^4x^4} \Big\{ \left[1 + 2g^2(5+2x^2) - g^4(23+12x^2) \right] K_1(2x) + x \left[1 + 10g^2 - g^4(23+4x^2) \right] K_0(2x) \\ &+ \frac{g^2m\pi e^{-2x}}{4Mx} \left[2(3g^2-2) \left(6x^{-1} + 12 + 10x + 4x^2 + x^3 \right) \right] + g^2x \left(2 + 4x + 2x^2 + 3x^2 \right) \Big\}, \\ V_S^{TPE}(r) &= \frac{g^4m^5}{32\pi^3f^4} \Big\{ 3xK_0(2x) + (3+2x^2)K_1(2x) - \frac{3\pi m e^{-2x}}{16Mx} \left(6x^{-1} + 12 + 11x + 6x^2 + 2x^3 \right) \Big\}, \\ W_S^{TPE}(r) &= \frac{g^2m^6}{48\pi^2f^4} \frac{e^{-2x}}{x^6} \Big\{ \left(c_4 + \frac{1}{4M} \right) \left(1 + x \right) \left(3 + 3x + 2x^2 \right) - \frac{g^2}{16M} \left(18 + 36x + 31x^2 + 14x^3 + 2x^4 \right) \Big\}, \\ V_{LS}^{TPE}(r) &= -\frac{3g^4m^6}{64\pi^2Mf^4} \frac{e^{-2x}}{x^6} \left(1 + x \right) \left(2 + 2x + x^2 \right), \\ W_{LS}^{TPE}(r) &= \frac{g^2(g^2 - 1)m^6}{32\pi^2Mf^4} \frac{e^{-2x}}{x^6} \left(1 + x \right)^2, \end{split}$$
Chiral 2π exchange (upto N²LO)

 $\mathcal{V}_{NN} \ = \ V_C(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_C(r) \ + \ \left[V_S(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_S(r) \right] \vec{\sigma}_1 \cdot \vec{\sigma}_2$

 $+ \left[V_T(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_T(r) \right] \left(3\vec{\sigma}_1 \cdot \hat{r}\vec{\sigma}_2 \cdot \hat{r} - \vec{\sigma}_1 \cdot \vec{\sigma}_2 \right) + \left[V_{LS}(r) + \vec{\tau}_1 \cdot \vec{\tau}_2 W_{LS}(r) \right] \vec{L} \cdot \vec{S} \,,$



Effects of the chiral 2π exchange are clearly visible in NN phases

Rentmeester et al.'99,'03 Birse, McGovern '06

Nucleon-nucleon scattering

Selected neutron-proton scattering observables at E_{lab} = 200 MeV (preliminary results with improved-chiral N³LO potential) 15 $d\sigma/d\Omega$ [mb/sr] D A_{v} 0.5 10 0.5 0 5 0000000 0000 -0.5 0.5 1 A_{XX} Α A_{yy} 0.5 0 0.5 0 0¢ -0.5 o o o o Nijmegen PWA -0.5 -0.5 N³LO -1 estimated uncertainty -1 60 120 180 60 120 180 60 120 0 180

At N³LO, 2N observables are accurately described up to at least $E_{lab} \sim 200 \text{ MeV}$

 θ_{CM} [deg]

0

 θ_{CM} [deg]

0

 θ_{CM} [deg]

Elastic Nd scattering with N³LO 2NF

Selected neutron-deuteron scattering observables at E_{lab} = 10 MeV (preliminary results with i_{mproved}-chiral N³LO potential)



Clear room for 3NF effects in A_y and iT₁₁

Notice: most of the data are Coulomb-corrected pd data

Elastic Nd scattering with N³LO 2NF

Selected neutron-deuteron scattering observables at E_{lab} = 70 MeV (preliminary results with i_{mproved}-chiral N³LO potential)



Clear room for 3NF effects in the cross section and A_{ij}

Elastic Nd scattering with N³LO 2NF

Selected neutron-deuteron scattering observables at E_{lab} = 135 MeV (preliminary results with i_{mproved}-chiral N³LO potential)



Clear room for 3NF effects

Three-nucleon force: Status and ongoing developments









3NF structure functions at large distance are model-independent and parameter-free predictions based on χ symmetry of QCD + exp. information on π N system



NLO (Q²)



NLO (Q²)



















The TPE 3NF has the form (modulo 1/m-terms):

 $V_{2\pi} = \frac{\vec{\sigma}_1 \cdot \vec{q}_1 \, \vec{\sigma}_3 \cdot \vec{q}_3}{[q_1^2 + M_\pi^2] [q_3^2 + M_\pi^2]} \Big(\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \, \mathcal{A}(q_2) + \boldsymbol{\tau}_1 \times \boldsymbol{\tau}_3 \cdot \boldsymbol{\tau}_2 \, \vec{q}_1 \times \vec{q}_3 \cdot \vec{\sigma}_2 \, \mathcal{B}(q_2) \Big)$



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N²LO [Q³]:
$$\mathcal{A}^{(3)}(q_2) = \frac{g_A^2}{8F_\pi^4} \left((2c_3 - 4c_1)M_\pi^2 + c_3q_2^2 \right), \qquad \mathcal{B}^{(3)}(q_2) = \frac{g_A^2c_4}{8F_\pi^4} \left(\frac{g_A^2}{8F_\pi^4} + \frac{g_A^2}{8F_\pi^4} \right)$$



The TPE 3NF has the form (modulo 1/m-terms):

N²LO [Q³]: van Kolck '94

$$V_{2\pi} = \frac{\vec{\sigma}_1 \cdot \vec{q}_1 \, \vec{\sigma}_3 \cdot \vec{q}_3}{\left[q_1^2 + M_\pi^2\right] \left[q_3^2 + M_\pi^2\right]} \left(\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \, \mathcal{A}(q_2) + \boldsymbol{\tau}_1 \times \boldsymbol{\tau}_3 \cdot \boldsymbol{\tau}_2 \, \vec{q}_1 \times \vec{q}_3 \cdot \vec{\sigma}_2 \, \mathcal{B}(q_2)\right)$$

$$\mathcal{A}^{(3)}(q_2) = \frac{g_A^2}{8F_\pi^4} \left((2\mathbf{c_3} - 4\mathbf{c_1})M_\pi^2 + \mathbf{c_3}q_2^2 \right), \qquad \mathcal{B}^{(3)}(q_2) = \frac{g_A^2 \mathbf{c_4}}{8F_\pi^4}$$



• N³LO [Q⁴]:
$$\mathcal{A}^{(4)}(q_2) = \frac{g_A^4}{256\pi F_\pi^6} \Big[A(q_2) \left(2M_\pi^4 + 5M_\pi^2 q_2^2 + 2q_2^4 \right) + \left(4g_A^2 + 1 \right) M_\pi^3 + 2 \left(g_A^2 + 1 \right) M_\pi q_2^2 \Big],$$
$$\mathcal{B}^{(4)}(q_2) = -\frac{g_A^4}{256\pi F_\pi^6} \Big[A(q_2) \left(4M_\pi^2 + q_2^2 \right) + (2g_A^2 + 1)M_\pi \Big] \qquad \text{Ishikawa, Robilotta '07} \\ \text{Bernard, EE, Krebs, Meißner '08} \Big]$$

The TPE 3NF has the form (modulo 1/m-terms):

 $V_{2\pi} = \frac{\vec{\sigma}_1 \cdot \vec{q}_1 \, \vec{\sigma}_3 \cdot \vec{q}_3}{[q_1^2 + M_\pi^2] [q_3^2 + M_\pi^2]} \Big(\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \, \mathcal{A}(q_2) + \boldsymbol{\tau}_1 \times \boldsymbol{\tau}_3 \cdot \boldsymbol{\tau}_2 \, \vec{q}_1 \times \vec{q}_3 \cdot \vec{\sigma}_2 \, \mathcal{B}(q_2) \Big)$

$$\mathcal{A}^{(3)}(q_2) = \frac{g_A^2}{8F_\pi^4} \left((2c_3 - 4c_1)M_\pi^2 + c_3 q_2^2 \right), \qquad \mathcal{B}^{(3)}(q_2) = \frac{g_A^2 c_4}{8F_\pi^4}$$



$$N^{3}LO [Q^{4}]: \qquad \mathcal{A}^{(4)}(q_{2}) = \frac{g_{A}^{4}}{256\pi F_{\pi}^{6}} \Big[A(q_{2}) \left(2M_{\pi}^{4} + 5M_{\pi}^{2}q_{2}^{2} + 2q_{2}^{4} \right) + \left(4g_{A}^{2} + 1 \right) M_{\pi}^{3} + 2 \left(g_{A}^{2} + 1 \right) M_{\pi}q_{2}^{2} \Big], \\ \mathcal{B}^{(4)}(q_{2}) = -\frac{g_{A}^{4}}{256\pi F_{\pi}^{6}} \Big[A(q_{2}) \left(4M_{\pi}^{2} + q_{2}^{2} \right) + \left(2g_{A}^{2} + 1 \right) M_{\pi} \Big] \qquad \text{Ishikawa, Robilotta '07} \\ \text{Bernard, EE, Krebs, Meißner '08}$$

 N⁴LO [Q⁵]: Krebs, Gasparyan, EE '12

N²LO [Q³]:
van Kolck '94

$$\begin{aligned} \mathcal{A}^{(5)}(q_{2}) &= \frac{g_{A}}{4608\pi^{2}F_{\pi}^{6}} \Big[M_{\pi}^{2}q_{2}^{2}(F_{\pi}^{2}\left(2304\pi^{2}g_{A}(4\bar{e}_{14}+2\bar{e}_{19}-\bar{e}_{22}-\bar{e}_{36})-2304\pi^{2}\bar{d}_{18}c_{3}\right) \\ &+ g_{A}(144c_{1}-53c_{2}-90c_{3})) + M_{\pi}^{4}\left(F_{\pi}^{2}\left(4608\pi^{2}\bar{d}_{18}(2c_{1}-c_{3})+4608\pi^{2}g_{A}(2\bar{e}_{14}+2\bar{e}_{19}-\bar{e}_{36}-4\bar{e}_{38})\right) \\ &+ g_{A}\left(72\left(64\pi^{2}\bar{l}_{3}+1\right)c_{1}-24c_{2}-36c_{3}\right)\right) + q_{2}^{4}\left(2304\pi^{2}\bar{e}_{14}F_{\pi}^{2}g_{A}-2g_{A}(5c_{2}+18c_{3})\right)\Big] \\ &- \frac{g_{A}^{2}}{768\pi^{2}F_{\pi}^{6}}L(q_{2})\left(M_{\pi}^{2}+2q_{2}^{2}\right)\left(4M_{\pi}^{2}(6c_{1}-c_{2}-3c_{3})+q_{2}^{2}(-c_{2}-6c_{3})\right), \\ \mathcal{B}^{(5)}(q_{2}) &= -\frac{g_{A}}{2304\pi^{2}F_{\pi}^{6}}\Big[M_{\pi}^{2}\left(F_{\pi}^{2}\left(1152\pi^{2}\bar{d}_{18}c_{4}-1152\pi^{2}g_{A}(2\bar{e}_{17}+2\bar{e}_{21}-\bar{e}_{37})\right)+108g_{A}^{3}c_{4}+24g_{A}c_{4}\right) \\ &+ q_{2}^{2}\left(5g_{A}c_{4}-1152\pi^{2}\bar{e}_{17}F_{\pi}^{2}g_{A}\right)\Big] + \frac{g_{A}^{2}c_{4}}{384\pi^{2}F_{\pi}^{6}}L(q_{2})\left(4M_{\pi}^{2}+q_{2}^{2}\right) \end{aligned}$$

Krebs, Gasparyan, EE '12

π N phase shifts in HB ChPT up to Q⁴ (KH PWA)



The determined values of LECs

	c_1	c_2	c_3	c_4	$\bar{d}_1 + \bar{d}_2$	$ar{d}_3$	\bar{d}_5	$\bar{d}_{14} - \bar{d}_{15}$	\bar{e}_{14}	$ar{e}_{15}$	\bar{e}_{16}	\bar{e}_{17}	\bar{e}_{18}
Q^4 fit to GW	-1.13	3.69	-5.51	3.71	5.57	-5.35	0.02	-10.26	1.75	-5.80	1.76	-0.58	0.96
Q^4 fit to KH	-0.75	3.49	-4.77	3.34	6.21	-6.83	0.78	-12.02	1.52	-10.41	6.08	-0.37	3.26

Krebs, Gasparyan, EE '12



The determined values of LECs

	c_1	c_2	<i>C</i> 3	c_4	$\bar{d}_1 + \bar{d}_2$	\bar{d}_3	\bar{d}_5	$\left \bar{d}_{14}-\bar{d}_{15}\right $	\bar{e}_{14}	$ar{e}_{15}$	\bar{e}_{16}	\bar{e}_{17}	\bar{e}_{18}
Q^4 fit to GW	-1.13	3.69	-5.51	3.71	5.57	-5.35	0.02	-10.26	1.75	-5.80	1.76	-0.58	0.96
Q^4 fit to KH	-0.75	3.49	-4.77	3.34	6.21	-6.83	0.78	-12.02	1.52	-10.41	6.08	-0.37	3.26

Long-range 3NF up to N⁴LO (preliminary)

EE, Gasparyan, Krebs, Schat, in preparation



Chiral expansion of the 3NF



Chiral expansion of the 3NF



...

Chiral expansion of the 3NF



Pion-nucleon system in Δ-full EFT up to Q⁴

Krebs, Gasparyan, EE, to appear

π N phase shifts in HB ChPT up to Q⁴ (KH PWA)



LECs from pion-nucleon scattering (HB ChPT) in units of GeV⁻ⁿ (fit to KH PWA)

	c_1	c_2	c_3	c_4	$\bar{d}_1 + \bar{d}_2$	\bar{d}_3	\bar{d}_5	$\bar{d}_{14} - \bar{d}_{15}$	\bar{e}_{14}	\bar{e}_{15}	\bar{e}_{16}	\bar{e}_{17}	\bar{e}_{18}
Δ-less approach	-0.75	3.49	-4.77	3.34	6.21	-6.83	0.78	-12.02	1.52	-10.41	6.08	-0.37	3.26
∆-full approach	-0.95	1.90	-1.78	1.50	2.40	-3.87	1.21	-5.25	-0.24	-6.35	2.34	-0.39	2.81
∆-contribution	0	2.81	-2.81	1.40	2.39	-2.39	0	-4.77	1.87	-4.15	4.15	-0.17	1.32

2π-exchange 3NF: Δ-full vs Δ-less EFT

Krebs, Gasparyan, EE, to appear



- Δ -full and Δ -less EFT predictions agree well with each other
- Δ-full approach shows clearly a superior convergence
- remarkably, the final 2π 3NF turns out to be rather weak at large distances...

2π -exchange 3NF at N⁴LO in Δ -less EFT

Krebs, Gasparyan, EE, to appear

Chiral expansion of TPE "structure functions" F_i (in MeV) in the equilateral-triangle configuration (Δ -less EFT)



For intermediate-range topologies, the effects of the Δ appear to be more pronounced (work in progress)

Partial wave decomposition of the 3NF

Low Energy Nuclear Physics International Collaboration (LENPIC) Bochum-Bonn-Cracow-Darmstadt-Iowa-Jülich-Kyushu-Ohio-Orsay

Faddeev equation is solved in the partial wave basis: $|p,q,\alpha\rangle \equiv |pq(ls)j(\lambda \frac{1}{2})I(jI)JM_J\rangle |(t\frac{1}{2})TM_T\rangle$

Too many terms for doing PWD "manually" → let computer do the job… Golak et al. EPJA 43 (2010) 241



→ feasible task but requires a few 10 MCPU hours for the N³LO 3NF...

It is possible to reduce the number of integrations to 2 exploiting locality of the 3NF Krebs, Hebeler

→ no need for supercomputers!

Current status

- PW matrix elements of the 3NF without regulator and with the old nonlocal regulator are available
- PWD of r-space regularized 3NF consistent with the new **i**chiral-**potential** in progress

Summary and outlook

Improved chiral NN potential up to N³LO

- Better performance at high energies, no fine tuning in π N LECs, no need for additional spectral function regularization, careful error estimation...
- Application to elastic Nd scattering shows clearly the need for 3NF (most striking at energies of E_{lab} ~ 70...150 MeV)

Chiral three-nucleon force

- Worked out completely at N³LO and at N⁴LO for 2π , 2π - 1π and ring graphs. The N⁴LO contributions are driven by the Δ and are large (as expected).
- Alternatively, calculations in EFT with explicit Δ are being performed. For 2π 3NF, both approaches lead to comparable results (with the Δ -full approach showing superior convergence). Δ -contributions to 2π - 1π and ring topologies have also been worked out, short-range terms in progress...
- Very good progress on the PWD of the 3NF

Future plans: completing derivation of the 3NF at N⁴LO and Δ contributions at N³LO; PWD of the locally regularized 3NF; 3NF effects in 3N scattering and spectra of light nuclei...

Two nucleons à la Weinberg

How to renormalize the Schrödinger equation Lepage, nucl-th/9697929

- 1. Introduce a *finite* cutoff $M_{\pi} \ll \Lambda \sim \Lambda_{hard}$ All symmetries can be preserved Slavnov '71; Djukanovic et al.'05, Hall, Pascalutsa '12
- 2. Tune $C_i(\Lambda)$ to low-energy observables \langle (implicit) renormalization
- 3. Check self-consistency by means of error-plots (Lepage-plots)

 $T = \frac{\alpha_1 + \alpha_2 \Lambda + \alpha_3 \Lambda^2}{\beta_1 + \beta_2 \Lambda + \beta_2 \Lambda^2}$

Predictive power easily understood in terms of Modified Effective Range Theory...

How not to renormalize the Schrödinger equation: an infinite cutoff limit

Removing Λ by taking the limit $\Lambda \to \infty$ may yield finite results for the amplitude but does not qualify for a consistent renormalization in the EFT sense. It is only justified if all necessary counterterms are included... EE, Gegelia, EPJA 41 (2009) 341

$$\begin{cases} \xrightarrow{\Lambda \to \infty} & T = \frac{\alpha_3}{\beta_3} \\ \text{renormalization} \\ \xrightarrow{} & T = \frac{\alpha_1 + \alpha_2 \mu + \alpha_3 \mu^2}{\beta_1 + \beta_2 \mu + \beta_3 \mu^2} \end{cases}$$

The cutoff issue

Why cutoff?

 $T = V + VG_0T = V + VG_0V + VG_0VG_0V + \dots$



increasingly UV divergent integrals are generated through iterations



Ideally, would like to calculate the integrals

$$T^{(n)} = V[G_0V]^{n-1} = m^{n-1} \int_0^{\Lambda} d^3 l_1 \dots d^3 l_{n-1} \frac{V(\vec{p}', \vec{l}_1) V(\vec{l}_1, \vec{l}_2) \dots V(\vec{l}_{n-1}, \vec{p})}{[p^2 - l_1^2 + i\epsilon] \dots [p^2 - l_{n-1}^2 + i\epsilon]}$$

subtract the UV divergences ($\sim \log \Lambda$; Λ ; Λ^2 ;...) and take the limit $\Lambda \to \infty$. This is not possible in practice (except for pionless EFT) \longrightarrow let Λ finite and adjust bare C_i to exp data (= implicit renormalization). Λ should be not taken too high [Lepage, EE, Meißner, Gegelia], in practice $\max[\Lambda] \sim 600$ MeV (otherwise spurious BS...)

Price to pay: finite cutoff artifacts (i.e. terms ~ $1/\Lambda$; $1/\Lambda^2$; $1/\Lambda^3$;...), may become an issue at higher energies (e.g. $E_{lab} \sim 200$ MeV corresponds to p ~ 310 MeV/c)

Is it possible to eliminate or at least reduce finite cutoff artifacts?